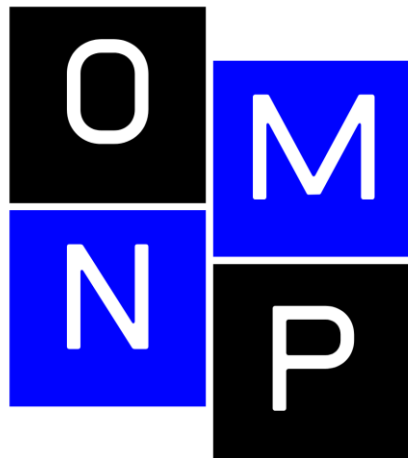


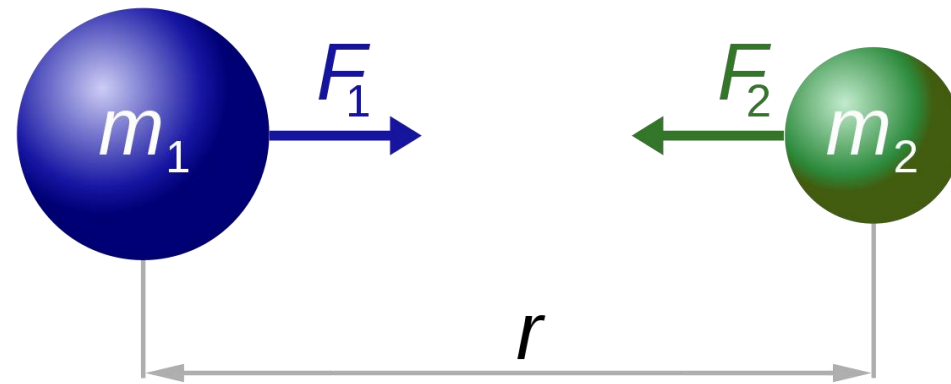
<https://youtu.be/vq0EoPcWWmE>

4 – A LEI DA GRAVIDADE

17 EQUAÇÕES QUE MUDARAM O MUNDO – OUTSPOKEN MARKET NA PRÁTICA – LEANDRO GUERRA



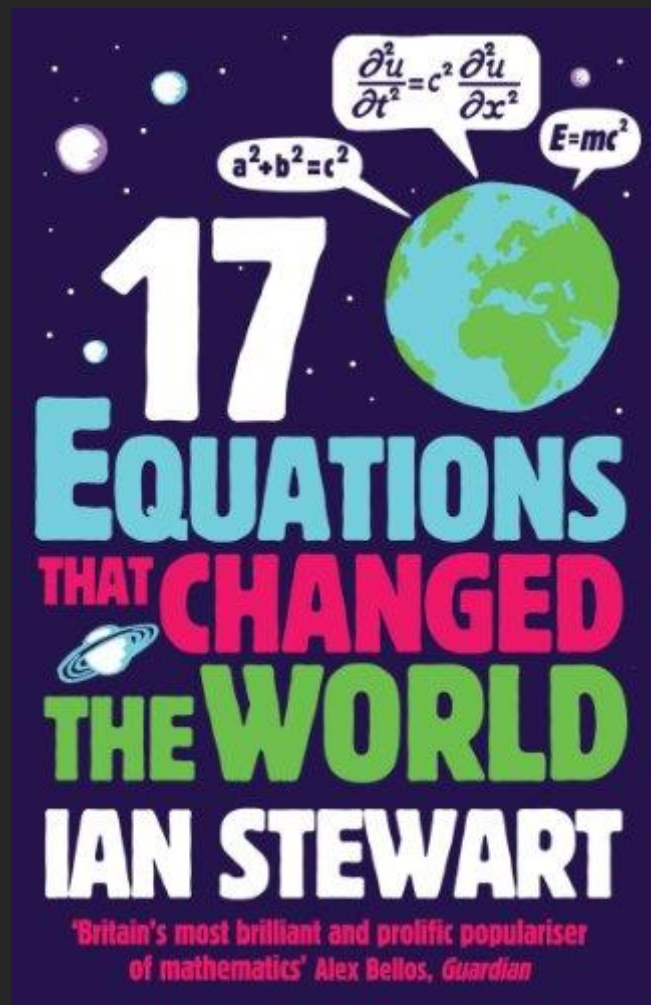
OUTSPOKEN MARKET
NA PRÁTICA



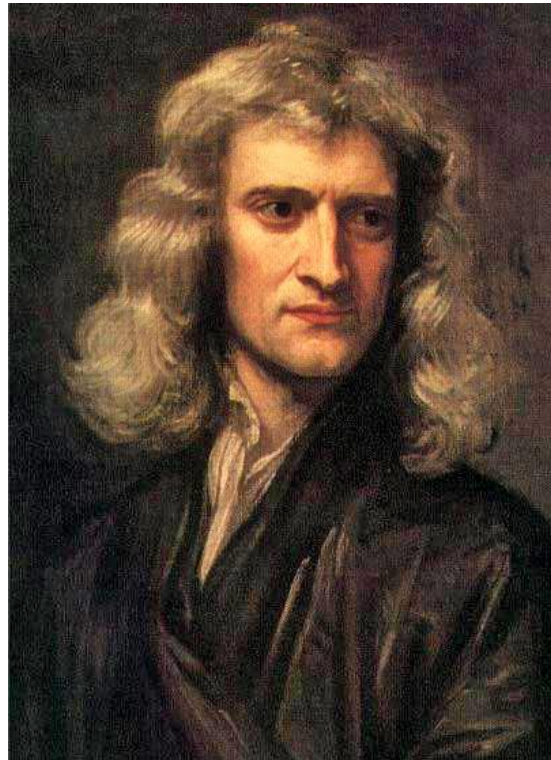
$$F_1 = F_2 = G \frac{m_1 \times m_2}{r^2}$$

17 Equations That Changed the World by Ian Stewart

- Pythagoras's Theorem** $a^2 + b^2 = c^2$ Pythagoras, 530 BC
- Logarithms** $\log xy = \log x + \log y$ John Napier, 1610
- Calculus** $\frac{df}{dt} = \lim_{h \rightarrow 0} \frac{f(t+h) - f(t)}{h}$ Newton, 1668
- Law of Gravity** $F = G \frac{m_1 m_2}{r^2}$ Newton, 1687
- The Square Root of Minus One** $i^2 = -1$ Euler, 1750
- Euler's Formula for Polyhedra** $V - E + F = 2$ Euler, 1751
- Normal Distribution** $\Phi(x) = \frac{1}{\sqrt{2\pi\rho}} e^{-\frac{(x-\mu)^2}{2\rho}}$ C.F. Gauss, 1810
- Wave Equation** $\frac{\partial^2 u}{\partial t^2} = c^2 \frac{\partial^2 u}{\partial x^2}$ J. d'Alembert, 1746
- Fourier Transform** $f(\omega) = \int_{-\infty}^{\infty} f(x) e^{-2\pi i x \omega} dx$ J. Fourier, 1822
- Navier-Stokes Equation** $\rho \left(\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} \right) = -\nabla p + \nabla \cdot \mathbf{T} + \mathbf{f}$ C. Navier, G. Stokes, 1845
- Maxwell's Equations** $\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0}$ $\nabla \cdot \mathbf{H} = 0$ $\nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{H}}{\partial t}$ $\nabla \times \mathbf{H} = \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t}$ J.C. Maxwell, 1865
- Second Law of Thermodynamics** $dS \geq 0$ L. Boltzmann, 1874
- Relativity** $E = mc^2$ Einstein, 1905
- Schrodinger's Equation** $i\hbar \frac{\partial}{\partial t} \Psi = H\Psi$ E. Schrodinger, 1927
- Information Theory** $H = -\sum p(x) \log p(x)$ C. Shannon, 1949
- Chaos Theory** $x_{i+1} = kx_i(1 - x_i)$ Robert May, 1975
- Black-Scholes Equation** $\frac{1}{2}\sigma^2 S^2 \frac{\partial^2 V}{\partial S^2} + rS \frac{\partial V}{\partial S} + \frac{\partial V}{\partial t} - rV = 0$ F. Black, M. Scholes, 1990



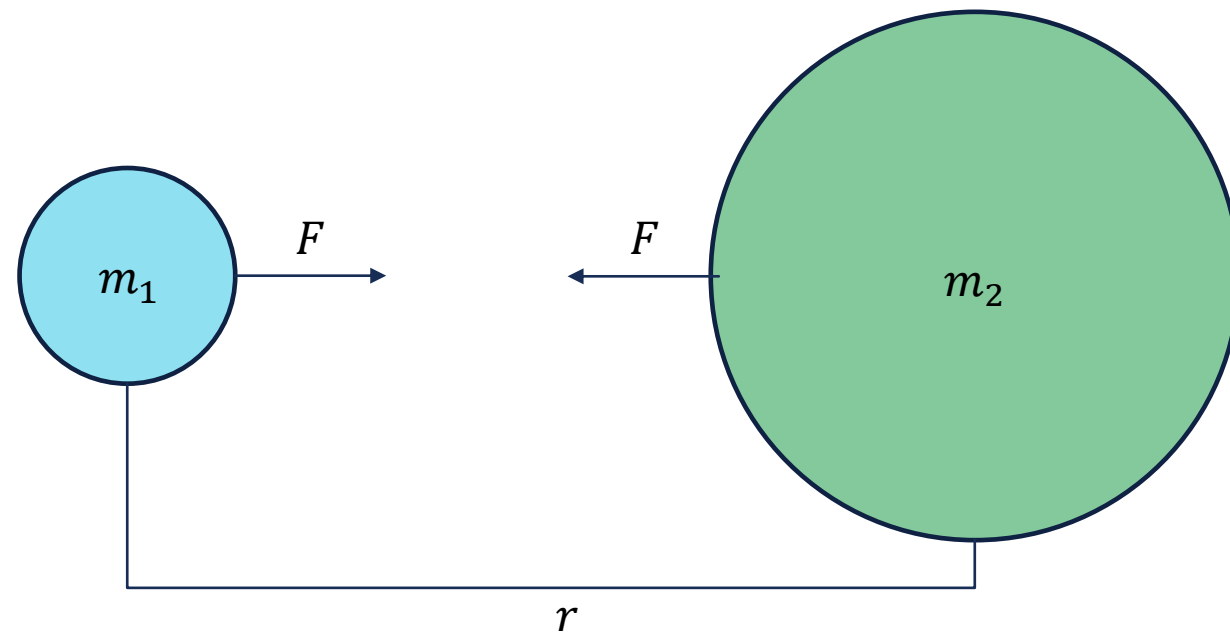
ISAAC NEWTON



Isaac Newton
(1642 -1726)

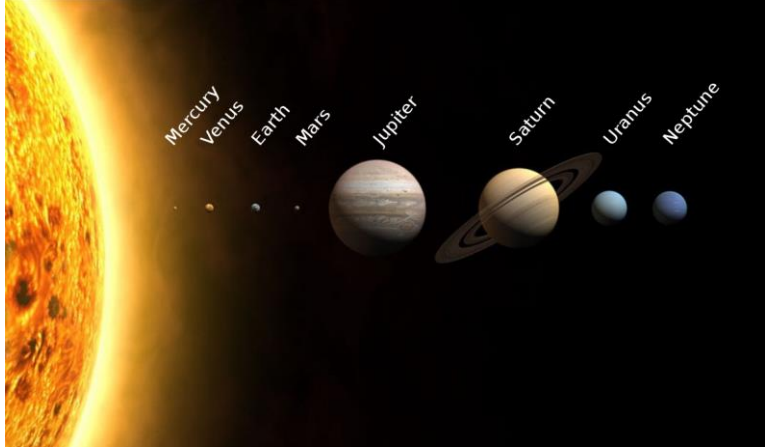
A FORÇA GRAVITACIONAL

$$F = G \frac{m_1 m_2}{r^2}$$



A teoria geral da relatividade de Einstein suplantou a lei geral da gravidade de Newton

POR QUE ELA É IMPORTANTE?



Mantém o sistema solar coeso



Mantém a nossa atmosfera por perto



Ajuda no fluxo da água nos rios e na precipitação de chuvas e neve.



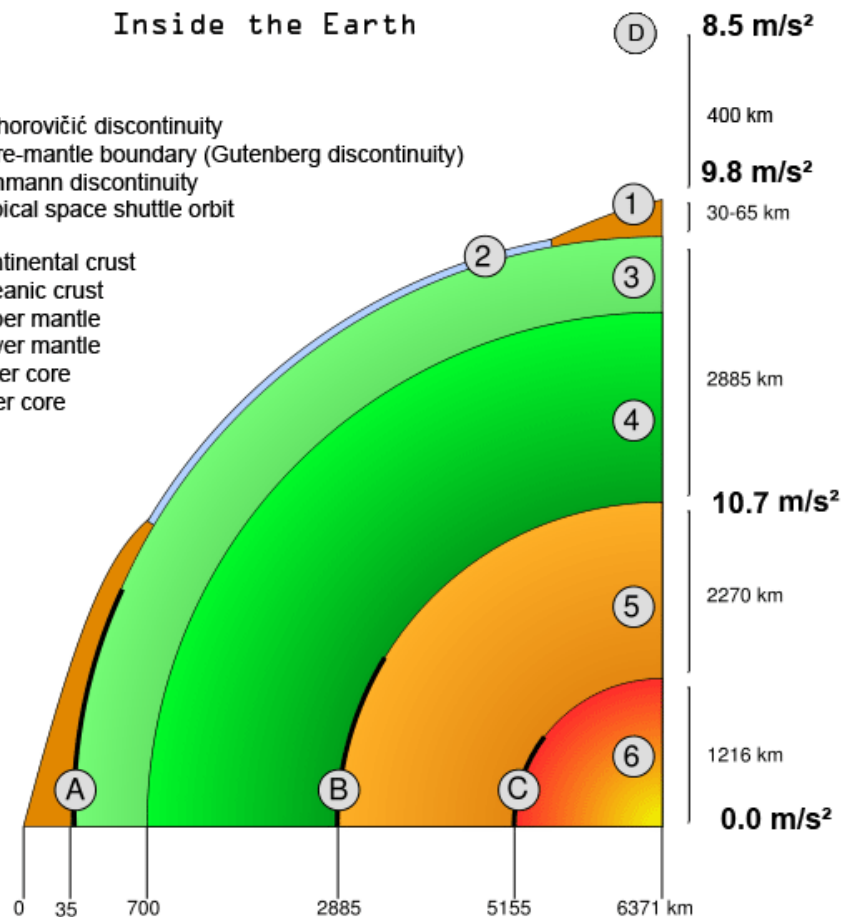
órbita da Lua e marés.

UMA OBSERVAÇÃO

Gravitational Field Strength: Inside the Earth

- A : Mohorovičić discontinuity
- B : Core-mantle boundary (Gutenberg discontinuity)
- C : Lehmann discontinuity
- D : Typical space shuttle orbit

- 1 : Continental crust
- 2 : Oceanic crust
- 3 : Upper mantle
- 4 : Lower mantle
- 5 : Outer core
- 6 : Inner core

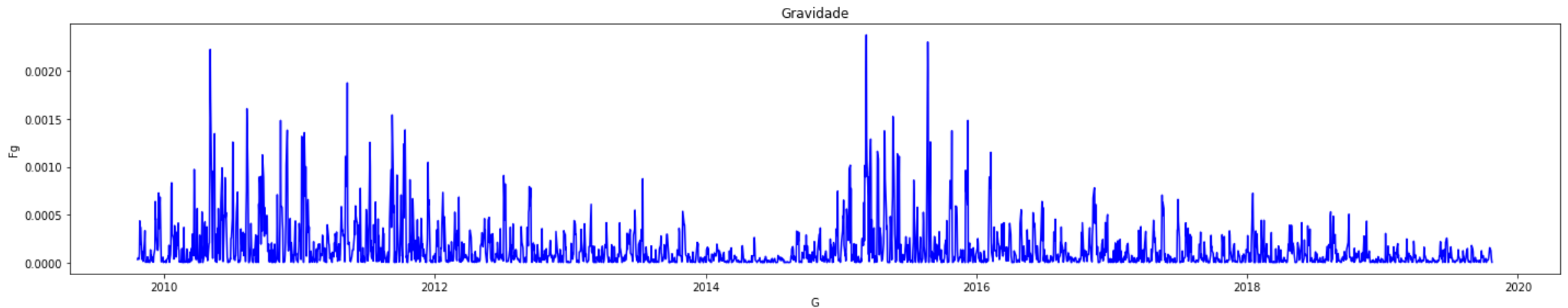


A força da gravidade esta sempre presente e é relativamente fraca quando comparada à outras forças.

Entretanto sua presença é notável quando tratamos de objetos com massas elevadas como os planetas ou outros corpos celestes.

APLICAÇÕES DA LEI DA GRAVIDADE

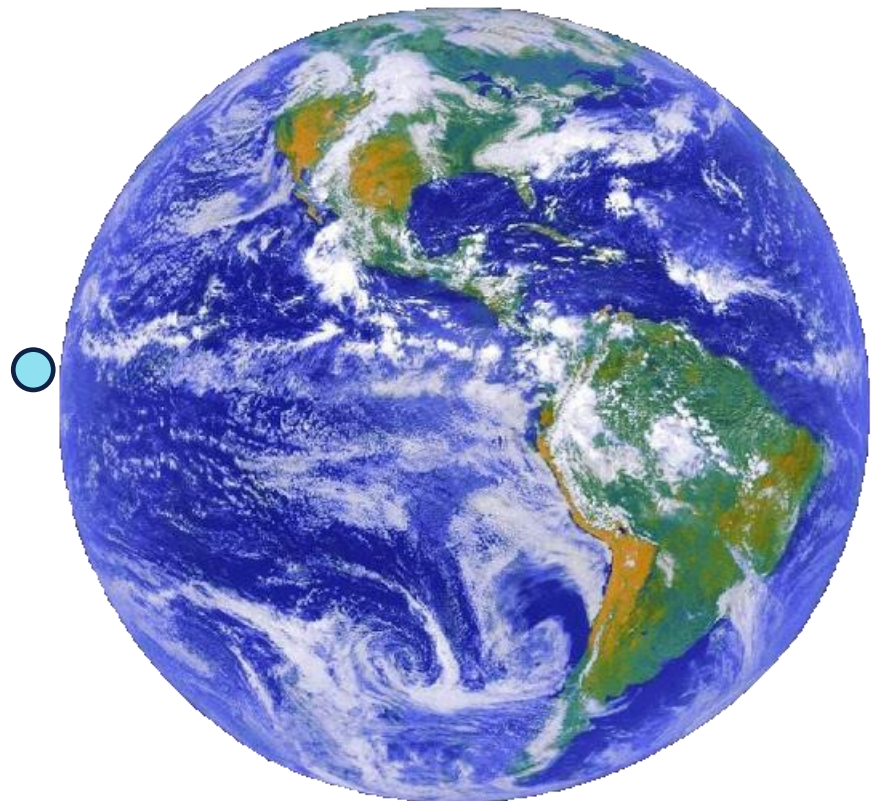
Variáveis preditivas em modelos do mercado financeiro



$$F = G \frac{m_1 m_2}{r^2}$$

APLICAÇÕES DA LEI DA GRAVIDADE

Cálculo razoavelmente aproximado da massa de objetos, como exemplo a massa da terra



$$F = G \frac{m_1 m_2}{r^2}$$

$$9,8 = 6,67 * 10^{-11} \frac{1m_2}{6371000^2}$$

$$9,8 * 4,06 * 10^{13} = 6,67 * 10^{-11} m_2$$

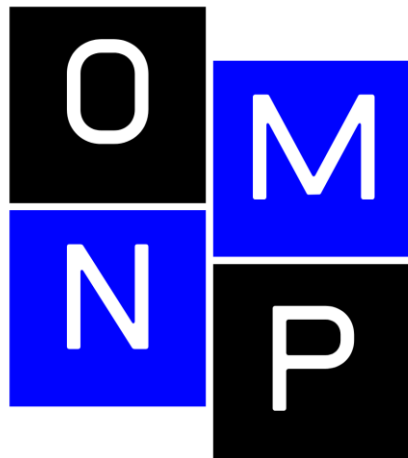
$$9,8 * 4,06 * 10^3 * \frac{1}{6,67} * 10^{11} = m_2$$

$$m_2 = 5.960.000.000.000.000.000.000$$

$$m_2 = 5,96 * 10^{24} kg$$

OBRIGADO E ATÉ A PRÓXIMA!

OUTSPOKEN MARKET



OUTSPOKEN MARKET
NA PRÁTICA

